

An Optimum Minimal Basis of Exponential Atomic Orbitals

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An analysis of some of the energetic properties of the conventional minimal STO basis is used to suggest a new optimum set of exponential functions for use in molecular calculations.

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In the famous paper of 1930 [1] proposing “Slater’s Rules” for the estimation of approximate self-consistent atomic orbitals, Slater made use of three pieces of information:

1) The asymptotic form for the solution of the Schrödinger equation for the hydrogen atom is

$$r^{n-1} \exp(-\zeta r)$$

for large values of r .

2) The idea of “screening” – replacing the potential generated by a “shell” of electrons by a spherical average.

3) The small amount of data on atomic calculations available at the time of writing.

Slater’s qualitative conclusions have remained unchallenged for nearly half a century – only quantitative changes have been made within his general framework:

a) For reasons of computational convenience the “quantum number” n^* of

$$r^{n^*-1} \exp(-\zeta r)$$

is now almost invariably an integer [2].

b) The orbital exponents ζ can now be accurately optimized – no approximate rules are required [3].

The minimal basis of STOs provides (either explicitly or implicitly) the background for most *ab initio* or semi-empirical calculations of molecular electronic structure and so it is worthwhile to subject this basis to a careful examination.

Taking points 1) to 3) above in reverse order, clearly b) has replaced 3) as a source of information. There can be no objection to 2) since it is the only way of defining “effective” analytical approximate AOs. However the argument for the use of STOs of form $r^{n-1} \exp(-\zeta r)$ because they are the asymptotic form of the hydrogenic Schrödinger equation does not, unfortunately, stand up to close examination. In the first place the asymptotic form (r^{n-1}) takes over from the exact form ($r^l L_{n+l}^{2l+1}(r)$) at rather large values of r ; there is, for example, still a 5 difference between the two $2s$ forms at 20 bohrs from the nucleus. Secondly, it is the *functional forms* $r^{n-1} \exp(-\zeta r)$ and $r^l L_{n+l}^{2l+1}(r) \exp(-\zeta r)$ which are asymptotically equivalent at large r : that is they both solve the radial differential equation at large r . But in order to use the approximate forms as *orbitals* they must be normalized, not in the asymptotic region, but in all space. Thus at $r=100$ bohrs the two $2s$ forms are essentially identical but normalization means that the corresponding orbitals differ by a factor of $\sqrt{3}$. It goes without saying of course that these enormous distances (20–100 bohrs) are of little consequence in the atomic energy or the interpretation of bonding. Classically, the electron at such large distances from the nucleus is moving extremely slowly. Indeed the kinetic energy value for an STO of quantum numbers n, l is always less than or equal to the hydrogenic value of $\frac{1}{2}\zeta^2$, being

$$\frac{n+2l(l+1)}{2n(n-1)}\zeta^2 \leq \frac{1}{2}\zeta^2$$

The average value of Z/r for the STO and hydrogenic orbitals are, however, the same; $-Z\zeta/n$. These considerations suggest that the use of a minimal basis of STOs may well lead to an imbalance between kinetic and potential energy and this analysis will be carried further using the virial theorem.

Even the most casual comparison of the results of the use of a minimal basis of STOs [3] for the calculation of atomic structure with Hartree–Fock results shows some important differences. The restricted functional form of the STOs prevents the minimal basis from reproducing the Hartree–Fock total energy, of course, but the most striking difference between the two sets of calculations is in the orbital energies. In particular the minimal STO basis for the heavier members of the first transition series (Fe–Zn) gives extremely poor orbital energies (see Fig. 2), zero or positive for Ni, Cu, Zn. Thus although the computed *total* energy for the minimal basis is correct to a fraction of one per cent, the orbital energies, which in a certain sense, give an indication of the distribution of the energy are in error by anything up to 90% in this important series.

The variational method of optimizing the STO exponents in the minimal basis calculation ensures that the virial theorem for the total energy is obeyed to a high

degree of accuracy by the resulting wave function (5-figure accuracy is typical). Thus if the total energy is a sum of kinetic (T), nuclear attraction (V_n) and electron repulsion (V_e) terms then

$$E = T + V_n + V_e \quad (1)$$

and

$$E = -T = \frac{1}{2}(V_n + V_e) \quad (2)$$

thus

$$E = (-T + V_e) - V_e \quad (3)$$

But if the orbital energies are ε_i and the orbital occupation numbers are n_i then

$$E = \sum n_i \varepsilon_i - V_e \quad (4)$$

Thus to a very good approximation the sum of the orbital energies is determined by the relative "weights" of the two essentially positive contributions to the total energy $-T + V_e$. Now examination of the results of the minimal STO basis calculation shows that the energy term $\sum n_i \varepsilon_i$ is usually smaller than the corresponding Hartree-Fock result, in particular it is much smaller for the transition series. Since the total energies are comparable, this implies that the electron repulsion energy V_e is too high in the approximate calculation; as we surmised, the kinetic and potential energy contributions are out of balance. This situation seems likely to be due to the use of the "higher" STOs which have low kinetic energy averages.

The lowest STOs of each atomic symmetry type are identical to the hydrogenic orbitals and therefore have kinetic energy $\frac{1}{2}\zeta^2$, the forms of the STOs of a given l value are all of the same general appearance except for the $1s$ function which differs from the other s functions in its behaviour at the origin. We have therefore used a minimal basis of STO functions consisting of only the $1s$, $2s$, $2p$, $3d$ forms to investigate the possibility of improving the energy distribution in atomic systems, while retaining, if possible, the very good total energy of the "usual" STO set. Preliminary exploratory work with the ground state of the sulphur atom was very encouraging - Table 1 gives the relevant data for comparison.¹ It is readily seen from Table 1 that the new minimal basis gives better agreement with the Hartree-Fock orbital energies in every case. The total energy is also very slightly lower but this is a rather trivial effect. Encouraged by these results we have performed optimizations on the elements from the second two rows of the periodic table since the results in [3] are already optimum for the first row. Table 2 gives the optimum orbital exponents and total energies for these atoms and Figs. 1 and 2 show the behaviour of the valence orbital energies compared to the Hartree-Fock and the conventional minimal basis.

There are two general points to be made about the results. Firstly there is an improvement in orbital energies in going from the conventional minimal basis to

¹ The last column of Table 1 verifies our conjecture about the unique nature of the $1s$ AO - total energy is poorer although there is some improvement in some orbital energies.

Table 1 Energy quantities for the ground state of the S atom

| Orbital energies | HF ^a | Best STO [3] | 1s, 2s, 2p | 1s, 2p |
|------------------|-----------------|--------------|------------|----------|
| 1s | -92.0046 | -91.8842 | -91.9346 | -91.9903 |
| 2s | -9.0045 | -8.7708 | -8.8201 | -8.4049 |
| 2p | -6.6827 | -6.4362 | -6.4843 | -6.5985 |
| 3s | -0.8796 | -0.8117 | -0.8517 | -0.7434 |
| 3p | -0.4374 | -0.3871 | -0.4335 | -0.4552 |
| Total Energy | -397.505 | -396.628 | -396.641 | -395.345 |

^a Numerical Hartree-Fock results are given in: Mann, J. B.: Technical Report LA-3690. University of California: Los Alamos Scientific Laboratory 1967, however the open-shell orbital energies differ from the analytical Roothaan-Hartree-Fock calculations given in Clementi, E.: Suppl. IBM J. Res. Develop. 9, 2 (1965) presumably due to the arbitrariness in the definition of the open-shell Fock operator (see: McWeeny, R.: Chem. Phys. Letters 35, 13 (1975)). The results reported here were obtained using the same formalism as the Clementi calculations, therefore in Table 1 and Figs. 1 and 2 "HF" refers to the Clementi calculations. This choice is for purposes of valid comparison and does not, of course, imply that Clementi's results obey a Koopmans' theorem or that Mann's do not.

Table 2. Optimized exponents for the 1s, 2s, 2p, 3d set^a

| Atom | 1s | 2s | "3s" | "4s" | 2p | "3p" | "4p" | 3d | Energy |
|------|---------|---------|--------|--------|---------|--------|--------|--------|------------|
| Na | 10.6232 | 3.2775 | 0.5636 | — | 3.4002 | — | — | — | -161.1225 |
| Mg | 11.5993 | 3.6803 | 0.7612 | — | 3.9112 | — | — | — | -198.8567 |
| Al | 12.5800 | 4.0829 | 0.9642 | — | 4.4861 | 0.9164 | — | — | -241.1533 |
| Si | 13.5618 | 4.4802 | 1.1601 | — | 4.9723 | 0.9770 | — | — | -288.0883 |
| P | 14.5435 | 4.8762 | 1.3470 | — | 5.4726 | 1.1226 | — | — | -339.9072 |
| S | 15.5253 | 5.2721 | 1.5309 | — | 5.9719 | 1.2655 | — | — | -396.6411 |
| Cl | 16.5066 | 5.6673 | 1.7095 | — | 6.4697 | 1.4184 | — | — | -458.5577 |
| Ar | 17.4881 | 6.0626 | 1.8834 | — | 6.9651 | 1.5767 | — | — | -525.8243 |
| K | 18.4688 | 6.4437 | 2.1026 | 0.5806 | 7.4629 | 1.8129 | — | — | -598.0820 |
| Ca | 19.4480 | 6.8113 | 2.3120 | 0.6322 | 7.9613 | 2.0508 | — | — | -675.5736 |
| Sc | 20.4300 | 7.2084 | 2.5118 | 0.6617 | 8.4627 | 2.2379 | — | 2.4621 | -758.3434 |
| Ti | 21.4114 | 7.6054 | 2.6791 | 0.6736 | 8.9620 | 3.3992 | — | 2.7796 | -846.7630 |
| V | 22.3939 | 8.0059 | 2.8524 | 0.6890 | 9.4606 | 2.5646 | — | 3.0484 | 1940.9306 |
| Cr | 23.3783 | 8.4086 | 3.0371 | 0.7006 | 9.9572 | 2.7275 | — | 3.2978 | -1040.9792 |
| Mn | 24.3602 | 8.8100 | 3.2000 | 0.7149 | 10.4556 | 2.8839 | — | 3.5504 | -1147.0974 |
| Fe | 25.3416 | 9.2072 | 3.3501 | 0.7333 | 10.9551 | 3.0434 | — | 3.7687 | -1259.0929 |
| Co | 26.3247 | 9.6087 | 3.5140 | 0.7465 | 11.4528 | 3.1989 | — | 3.9930 | -1377.4023 |
| Ni | 27.3075 | 10.0099 | 3.6726 | 0.7605 | 11.9494 | 3.3567 | — | 4.2157 | -1502.0994 |
| Cu | 28.2906 | 10.4120 | 3.8314 | 0.7734 | 12.4462 | 3.5105 | — | 4.4386 | -1633.3156 |
| Zn | 29.2737 | 10.8199 | 4.0033 | 0.8196 | 12.9410 | 3.6587 | — | 4.6612 | -1771.2540 |
| Ga | 30.2606 | 11.2021 | 4.1915 | 0.9751 | 13.4333 | 3.8535 | 0.7539 | 5.0569 | -1916.5290 |
| Ge | 31.2391 | 11.5593 | 4.3542 | 1.1215 | 13.9233 | 4.0499 | 0.8706 | 5.4412 | -2068.4435 |
| As | 32.2278 | 11.9574 | 4.5888 | 1.2878 | 14.4061 | 4.2471 | 0.9044 | 5.8128 | -2227.0823 |
| Se | 33.2068 | 12.3157 | 4.7764 | 1.4041 | 14.8916 | 4.4567 | 1.1248 | 6.1697 | -2392.4892 |
| Br | 34.1842 | 12.6747 | 4.9326 | 1.5608 | 15.3801 | 4.6528 | 1.2401 | 6.5305 | -2564.7889 |
| Kr | 35.1761 | 13.0791 | 5.1958 | 1.6421 | 15.8583 | 4.8447 | 1.3642 | 6.8815 | -2744.1078 |
| Atom | 1s | 2s | "3s" | "4s" | 2p | "3p" | "4p" | 3d | Energy |

^aThe columns labelled "3s", "4s" refer to 2s-type STO's; those labelled "3p", "4p" to 2p STO's.

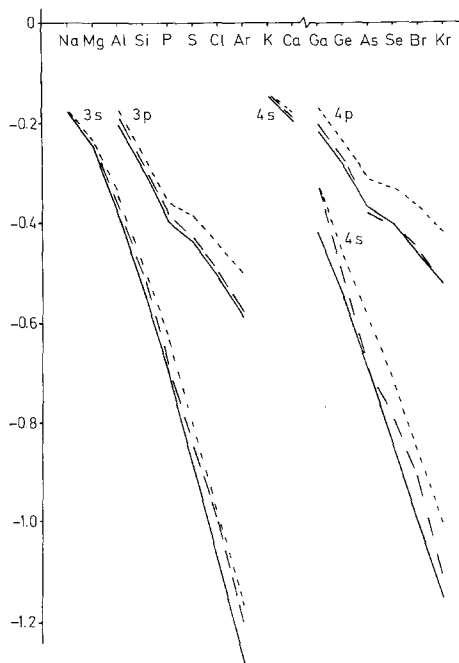


Fig. 1. Orbital energies for the “valence orbitals” of the elements Na to Kr (excluding the transition series). Full line Hartree-Fock (see Footnote to Table 1), broken line this work, dotted line results of Ref. [3]

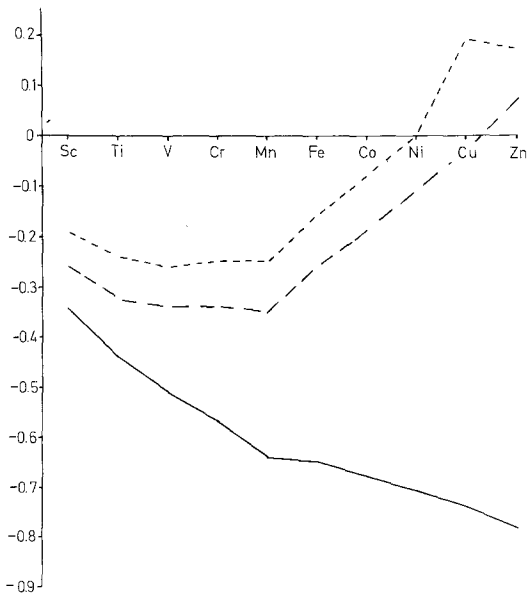


Fig. 2. Orbital energies of the 3d orbital in the first transition series (Sc to Zn). The lines are described in Fig. 1. (The 4s orbital energies are omitted as they are scarcely distinguishable on the scale used)

the $1s, 2s, 2p, 3d$ basis – this improvement being most dramatic in the case of the $3d$ energies of the first transition series. Secondly the total energies of the two minimal bases are essentially the same: sometimes the result computed here is the lower, sometimes the result of [3] is lower, the difference being very small indeed. Notwithstanding the first of these conclusions, the $3d$ orbital energies of the first transition series are, absolutely speaking, still in extremely poor agreement with the HF values as can be seen from Fig. 2. (It is also incidentally the case that the Hartree–Fock orbital energies for these systems are in very poor agreement with experiment – they are generally a factor of 2 too tightly bound in the Hartree–Fock model.) Thus the use of a minimal basis of $3d$ STOs for these atoms in a molecular calculation is almost guaranteed to give misleading results. Examination of the change in orbital energy with atomic number (for the minimal basis $3d$ energies) shows that even the wrong trends are given for changes in the population of the $3d$ levels. Preliminary work with the cobalt atom shows that the use of the $1s, 2s, 2p, 3d$ minimal basis plus one extra $3d$ function gives orbital energies within one or two per cent of the Hartree–Fock values for all orbitals with a total energy two hartrees better than the minimal $1s, 2s, 2p, 3d$ basis. The choice of optimum small bases of exponential functions for transition elements will be the subject of a later note.

Finally it is perhaps worth noting that the molecular energy integrals involving the “lower” STOs are rather simpler to compute than the orbitals of higher n ; so that there are considerable economies to be obtained by the use of the $1s, 2s, 2p, 3d$ basis. For example, existing $n=2$ programs for CNDO, INDO etc. can be used without further modification for second row atoms.

References

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